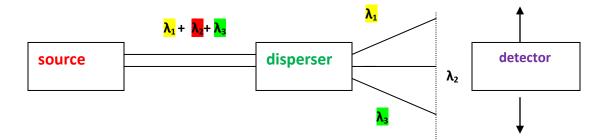
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Basics of spectroscopy

Any spectroscopic system consists essentially of 3 basic elements:



Spectroscopy

<u>Continuous radiation :</u> radiation consisting of a wide range of frequencies or wavelengths

<u>Monochromatic radiation :</u> (one colour)- radiation consisting of a narrow range of frequencies or wavelengths .

Source : Source of radiation under investigation .

Disperser : spreads or separates the light out spatially into its constituent wavelengths

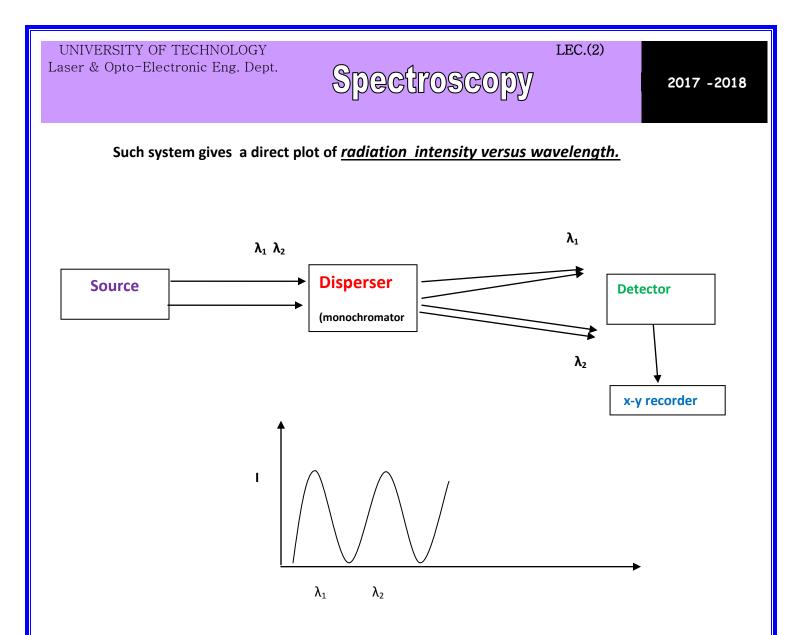
(or spectrum)

<u>Detector</u>: records the incident intensity of the separated monochromatic radiation as a function of position (ie wavelengths).

In practice , the disperser is scanned , and

the o/p signal from the detector is fed directly to the y - axis of a chart recorder (or pc),

and the wavelength signal is derived from the scanning disperser & fed to the x-axis of a chart recorder (or pc).



The Nature & Propagation of Light

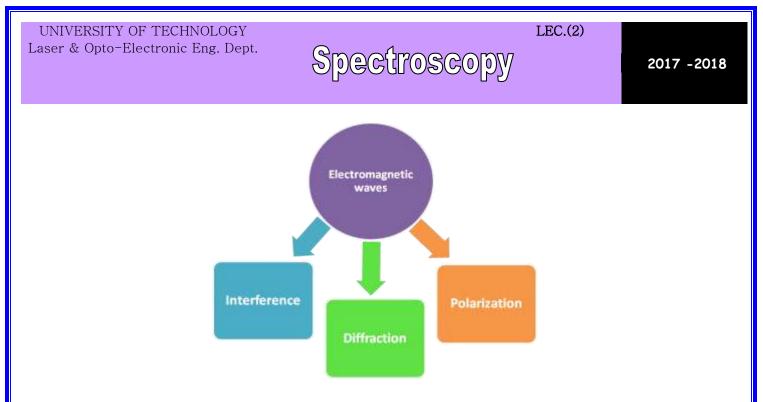
Light was thought to consist of a stream of some sort of *particles or corpuscles* emitting from a light source (the middle of the seventeenth century), (**Newton** and others).

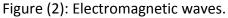
Huygens proposed that light is wave phenomena, (1678).

In the 19th century evidence of the wave theory of light grew up.

Fresnel and young experiments on interference and diffraction supported the wave theory.

Existence of electromagnetic waves ⇒ Hertz





Light is an electromagnetic wave \Rightarrow **Maxwell** (speed of propagation in 1873).

<u>Classical electromagnetic theory failed to explain several phenomena associated</u> <u>with absorption and emission of light.</u>

Example

Photoelectric emission (ejection of electrons when a light incidented on a conductor). **photoelectric effect**: by Hertz 1887

Which is a Liberation of electrons from the surface of a conductor when light strikes the surface.

Also <u>**x-ray production**</u> greater frequency than visible light & $\lambda \sim 1-20 \text{ A}^\circ$.

All these led to \Rightarrow **<u>quantum theory of radiation</u>** called **<u>photons</u>** or **<u>quanta</u>.**

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In 1905 Einstein extended Planck's idea that the energy in a light beam was concentrated in packets or <u>photons</u>.

Spectroscopy

The wave nature was retained that this photon is still has frequency, and its energy is proportional to the frequency E = hv

So, Propagation of light \Rightarrow waves.

Interaction of light with water \Rightarrow corpuscles (particles).

i.e, Dual nature of E.M. radiation.

Sources of light

All bodies emit electromagnetic radiation due to their molecules thermal motion (thermal radiation)

This radiation is a mixture of different wavelengths.

At 300°C the most intense of these waves is 5000 nm or 50000 Å (infrared region).

At 800°C the body emits visible (red *hot*).

At 3000°C the body appears white (emits 4000-7000 Å).

Examples:

- 1- (the tungsten lamp filament) Argon gas introduced to reduce evaporation of filament.
- 2- Carbon arc (very intense).

Two Carbon rods 10-20 cm long, 1 cm *dia* connected to 110-220 V dc, touched together momentarily and pulled apart few millimeters causes intense light due to electron bombardment of the positive rod (temperature 4000°C).

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- 3- Arc discharge in a metal vapor (Na, Hg).
- 4- Fluorescent light (a glass tube with argon and droplet of mercury (Hg)).

Spectroscopy

The amount of U.V. is absorbed by thin layer of *phosphor* (while coating on the glass wall).

5- LASER (narrow beam, monochromatic, very intense, many applications).

<u>Absorption & Emission of light</u> <u>Absorption</u>

No material is perfectly transparent. When light passing a medium, it partially absorbed increasing the internal energy in the material. The intensity (power per unit area) is correspondingly attended. When a beam of light of intensity (I) passes through a sheet of thickness (dx), the decrease dI in intensity is proportional to the initial intensity and thickness.

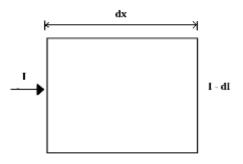


Figure 1): Transparent medium.

i.e. $dI \propto I \, dx \Longrightarrow dI = \propto I \, dx$

 $\propto \equiv$ absorption coefficient.

 $dI = -I \propto dx$

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$$\int_{I_0}^{I} \frac{dI}{I} = \int_0^x \propto dx$$
$$\ln \frac{I}{I_0} = -\infty x$$

for finite hickness x of the slab

Spectroscopy

$$I = I_0 e^{-\alpha x} \qquad Lambert's \ law$$

where I_0 is the intensity at x = 0.

 \propto is dependent on λ and polarization.

Example

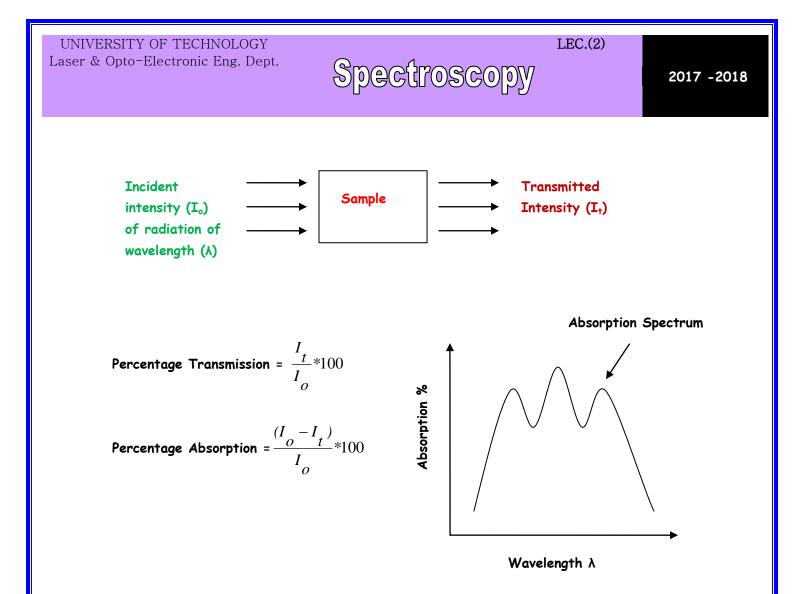
 $\propto = 10m^{-1}$ light passing a glass slab decrease by ^{5%}, what is the value of x ?

Solution:

$$\ln \frac{I}{I_0} = -\infty \ x \Longrightarrow$$
$$x = -\frac{1}{\infty} \ln \frac{I}{I_0} = -\frac{1}{10} \ln 0.95 = 0.5 \ cm.$$

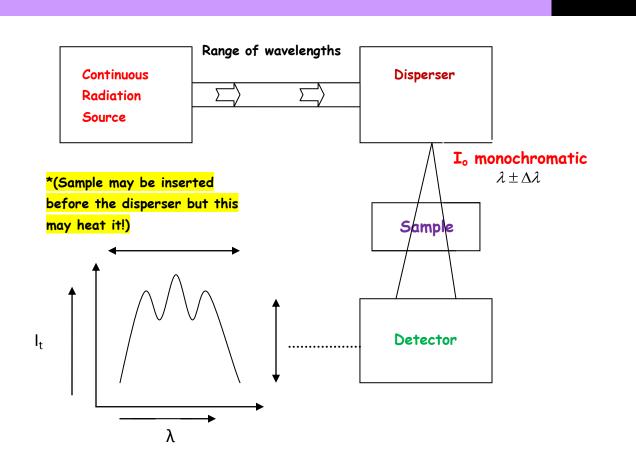
Absorption spectroscopy:

Absorption spectrum of a sample is the amount of radiation absorbed at each wavelength as a function of wavelength.



The apparatus required to produce such a spectrum is common to all regions of the spectrum, although the nature of components will vary for the different regions of the spectrum.

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Spectroscopy

Emission spectroscopy:-

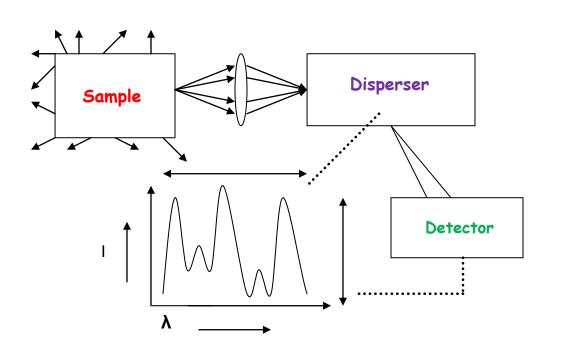
The emission spectrum of a sample is a display of the intensity of radiation at each wavelength.

*In emission spectroscopy the sample acts as the radiation source.

*The sample must be excited to emit radiation; e.g. gaseous samples may be excited by an electrical discharge.

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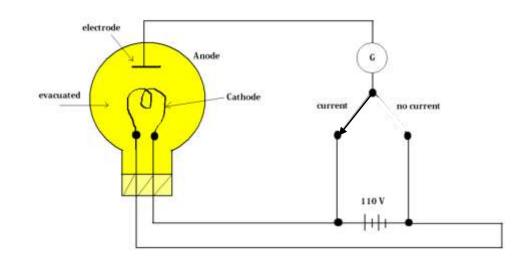
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Emission of line spectra \implies light emitted from atoms_heated in flame or excited electrically such as neon tubes \implies Contains **lines** <u>not</u> continuous wavelength \implies can be resolved by spectrometer (prism, diffraction grating, ...).

Spectroscopy

Thermionic emission discovered by Edison (1883)



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*Minimum energy required to escape from the surface is the <u>work function</u>, which varies for different materials.

Spectroscopy

Work function ~ few e.v.

 $1 e. v. = 1.6 \times 10^{-19} J$

Energy $\equiv U = qv$ = 1.6 × 10⁻¹⁹ coulomb × 1v 1 Ke. v. = 10^{3e.v.}

- $1 Me. v. = 10^{6e.v.}$
- $1 \text{ Ge. } v = 10^{9e.v.}$

Example

A particle having charge of 2e moves between 2 points of potential difference 1000 v, find P.E. in J and e.v.?

Solution:

$$U = qv = 2 \times 1.6 \times 10^{-19} \times 10^{3} = 3.2 \times 10^{-16} J$$

Then $\frac{3.2 \times 10^{-16}}{1.6 \times 10^{-19}} = 2000 \, e. \, v.$

At room temperature, no electrons escape.

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The Photoelectric Effect

If a metal is illuminated by light of **short** λ , electrons can have enough energy to **escape from a metal surface** even at law temperature (this is **P.E. effect**).

Spectroscopy

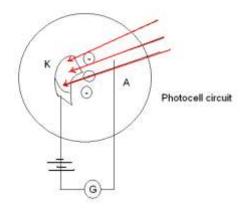


Figure (3): Photocell circuit.

The battery causes the electric field in the direction from (A) (Anode or collector) to (K) (cathode), and (e) (electron) emitted from (K) and went to (A).

(A) and (K) are enclosed in an evacuated container, photoelectric current is measured by the Galvanometer (G).

For a given cathode material (called emitter) there is minium frequency of light incidented on it to emit electrons called **Threshold frequency** which is for metals mostly in U.V., (λ_{emit} 200-300 nm) but for Potassium and Cesium oxide it is (400-700 nm).

To stop electrons, polarity reversed and potential energy e.v.> k.E. of electrons $(\frac{1}{2}mv_{max}^2)$.

To make direct measurement of k.E. of electron:

LEC.(2)

This electrical reversed potential is called *stopping potential* (V₀).

To make direct measurement of k.E. of electrons:

Explanation

By Einstein light quanta (photons) has energy

$$E = hf$$

Where
$$h \equiv Planck'sconstant = 6.63 \times 10^{-34}$$
 J.s

So the photon when hits a metal surface will give energy to the electron, so it will escape from the surface.

Spectroscopy

In leaving the surface, the (e) losses energy (ϕ) (the work function of the surface).

Combining with equation (1), get:

Example

If stopping potential is 3V for light wavelength 300nm, 2V for 400 nm and 1V for 600 nm determine the work function for this material and find the value of Planck's constant?

Solution:

According to equation (3):

$$eV_0 = hf - \varphi$$

Spectroscopy

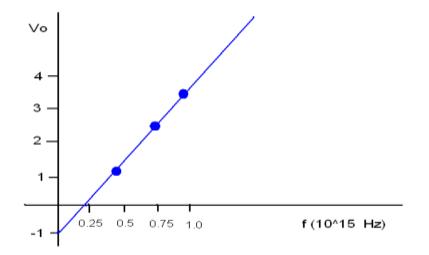
$$\therefore V_0 = \frac{h}{e}f - \varphi$$

$$\frac{h}{e} = slope = \frac{\Delta y}{\Delta x} = \frac{1}{0.25 \times 10^{15}}$$
Using $f = \frac{c}{\lambda}$

*The energy of photon E = hf

$$E = \frac{hc}{\lambda} \qquad since \quad f = \frac{c}{\lambda}$$

:
$$f = (0.5, 0.75, 1) \times 10^{15} s^{-1}$$





$$-\frac{\varphi}{e} \equiv intercept \ of \ graph = -1 \ v$$

$$\therefore \varphi = 1 e.v.$$

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Spectroscopy

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and
$$\frac{h}{e} = slope = \frac{\Delta y}{\Delta x} = \frac{1}{0.25 \times 10^{15}}$$

= $4 \times 10^{-15} \ J.C^{-1}.s$
 $\therefore \quad h = (4 \times 10^{-15}) \times 1.6 \times 10^{-19}$
= $6.4 \times 10^{-34} \ J.s$