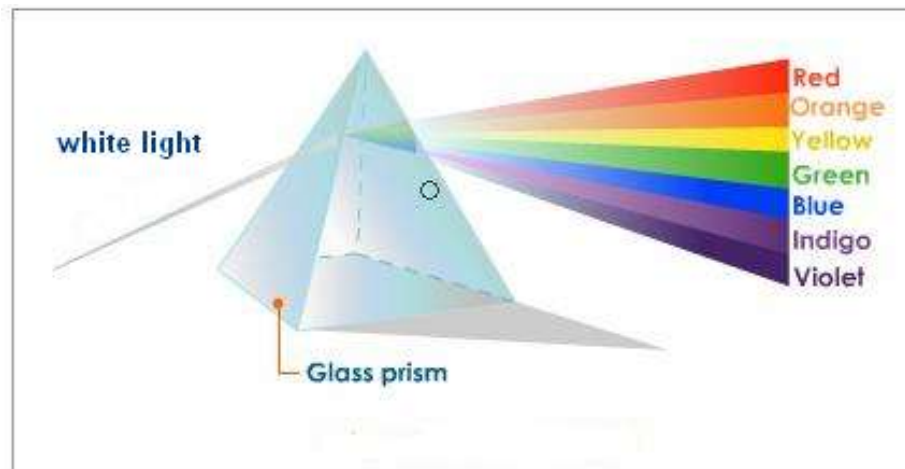


Line Spectra



The visible light spectrum.

White light \Rightarrow continuous spectrum (all wavelengths present).

Discharge in gas \Rightarrow few colors appear. (Isolated lines).

This is a line spectrum.

EXAMPLE

Discharge in :

Hydrogen

Sodium

Iron

Energy Levels

Every element has a characteristic line spectrum \Rightarrow result from the structure of atoms of the element.

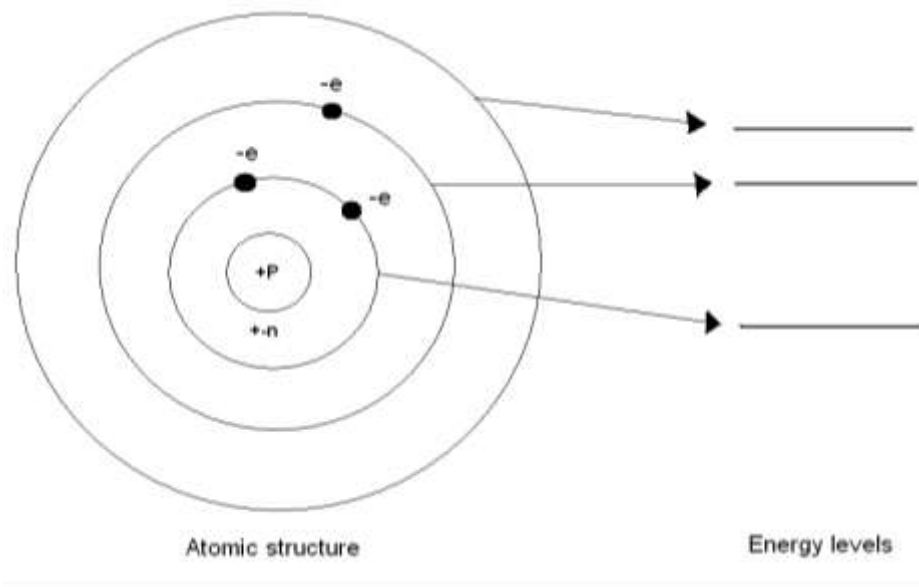


Figure (6): Atomic structure.

Bohr model of the atom:

The spectrum of H-atoms is explained by Bohr using 3 basic postulates:

1. The electron in H-atom can rotate about the nucleus in certain fixed orbit of radius (r), where orbital angular momentum L is a multiple of $\frac{h}{2\pi} (\hbar) \rightarrow$ i.e. Angular momentum is quantized.

$$L = I\omega = mvr = n\hbar = \frac{nh}{2\pi}$$

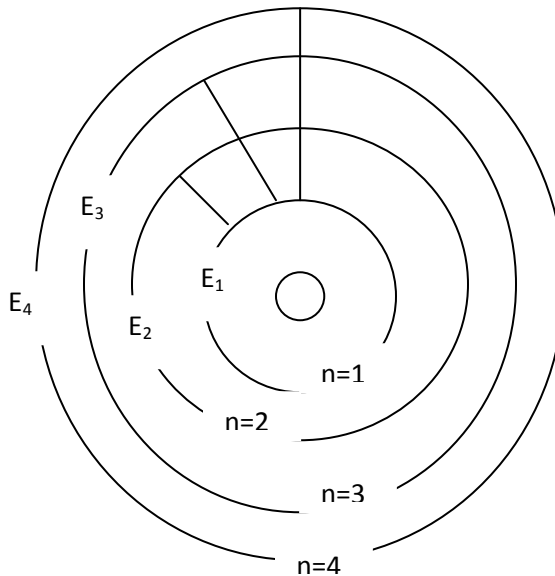
$$n = \text{integer } 1, 2, 3, 4, \dots, \infty$$

m = mass of the electron

v = linear velocity

2. The electron in the **stationary orbit (or state) does not emit electromagnetic radiation.**
3. **Radiation emitted or absorbed, When an electron undergoes a transition from one orbit to another, the energy of absorbed or emitted light photon is:**

$$\Delta E = E_1 - E_2 = h\nu$$



Now to derive the energy equation in order to understand the energy levels:

The electrostatic force = The centripetal force

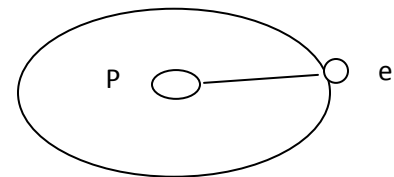
$$Kq_1q_2/r_2 = mv^2/r$$

Hence:

$$r = kze^2/mv^2 \quad \dots\dots\dots (1)$$

$$\text{Angular momentum} \quad L = I\omega = n\hbar \quad \dots\dots\dots (2)$$

$$\text{Since } I = mr^2 \quad , \quad \omega = v/r$$



This can be derived as follows:

$$\theta = s/r$$

$$d\theta = ds/r$$

$$d\theta/dt = (1/r) ds/dt$$

Hence: $\omega = v/r$

So:

$$mr^2v/r = n\hbar$$

$$v = n\hbar/mr \quad \dots\dots\dots(3)$$

Substitute (3) into (1), get:

$$r = n^2\hbar^2 / kze^2 m \quad \dots\dots\dots(4)$$

Substitute (4) into (3):

$$V = (n\hbar/m) (kze^2m/ n^2\hbar^2)$$

$$V = kze^2/n\hbar \quad \dots\dots\dots(5)$$

Total energy:

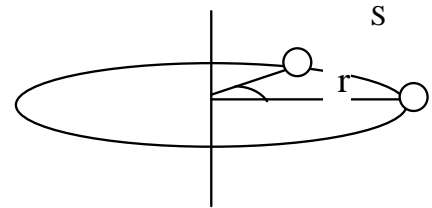
$$E_t = E_k + E_p$$

$$= 1/2mv^2 + E_p$$

$$\text{Work done } W = \int Fdr$$

$$= \int (kze^2/r^2) dr$$

$$= kze^2 \left[(1/r) \right]_{r_a}^{r_b}$$



$$= -kze^2/r_b + kze^2/r_a$$

If $r_b=r$, $r_a = \infty$

So :

$$W = -kze^2/r = E_p \quad \dots\dots\dots(6)$$

So:

$$E_t = 1/2mv^2 + (- kze^2 /r)$$

Substitute for the value of (v) from (5) :

$$E_t = - mz^2e^4k^2/2n^2\hbar^2$$

The negative sign is due to the connection between the nucleus and the electron.

For hydrogen atom

$$z = 1$$

$$E_t = \frac{[(-9.1 \times 10^{-31}) \times (1.6 \times 10^{-19})^4 \times (9 \times 10^9)^2]}{2 \times (6.63 \times 10^{-34} / 6.28)^2 n^2}$$

$$= -2179.6 \times 10^{-21} / n^2 \text{ Joule}$$

Hence:

$$E_t = -13.6/n^2 \text{ ev}$$

(Dividing by the charge of the electron).

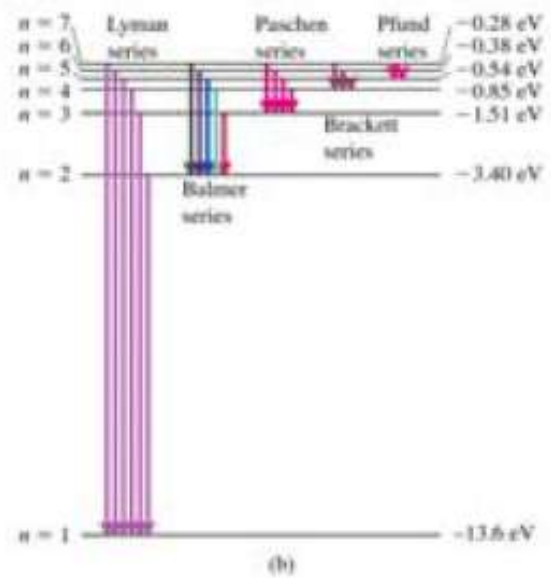
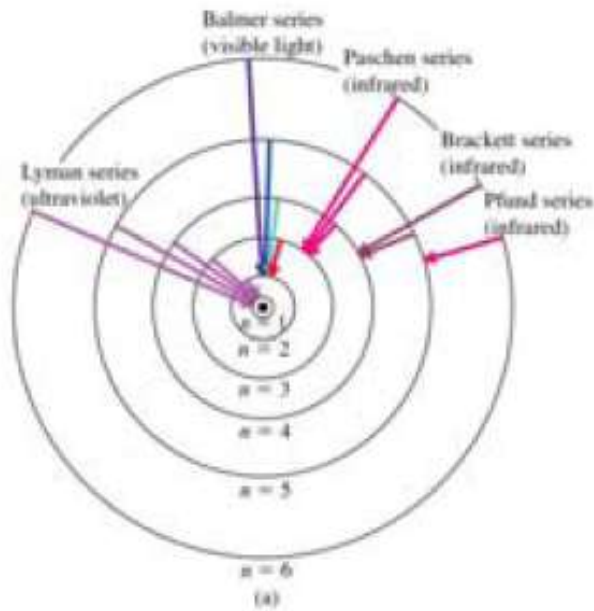
Now according to Bohr theory:

$$n=1 \quad E_1 = -13.6\text{eV}$$

$$n=2 \quad E_2 = -13.6/4 = -3.4\text{ eV}$$

$$n=3 \quad E_3 = -1.51\text{ eV}$$

$$n=\infty \quad E_\infty = 0$$



Now to calculate (λ) for the spectrum of the H-atom

$$E=hu = hc/\lambda$$

$$\lambda = hc/E$$

1st orbit For transition $1 \longrightarrow \infty$

$$\lambda_{(\infty-1)} = (6.63 \times 10^{-34} \times 3 \times 10^8) / (13.6 \times 1.6 \times 10^{-19}) = 914 \times 10^{-10} \text{ m}$$

$$= \mathbf{91.4 \text{ nm (u.v region)}}$$

For line $2 \longrightarrow 1$

$$E = 13.6 - 3.4 = 10.2 \text{ eV.}$$

$$\text{So } \lambda_{2 \rightarrow 1} = (6.63 \times 10^{-34} \times 3 \times 10^8) / (10.2 \times 1.6 \times 10^{-19}) = \mathbf{121.8 \text{ nm (u.v)}}$$

2nd orbit

$$\lambda_{(\infty-2)} = (6.63 \times 10^{-34} \times 3 \times 10^8) / (3.4 \times 1.6 \times 10^{-19}) = \mathbf{365.6 \text{ nm (u.v region)}}$$

$$\lambda_{(3-2)} = (6.63 \times 10^{-34} \times 3 \times 10^8) / (1.9 \times 1.6 \times 10^{-19}) = \mathbf{654.2 \text{ nm (visible)}}$$

(3.4-1.5)

Finding Line Wavelength (or Frequency)

Balmer Series

$$\frac{1}{\lambda} = R \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$$

$R \equiv$ Rydberg constant.

$n \equiv 3, 4, 5, \dots$

$R \equiv 1.097 \times 10^{-7} \text{ m}^{-1}$

$\lambda \equiv$ wavelength in m.

If $n = 3$

$$\frac{1}{\lambda} = 1.097 \times 10^7 \times \left(\frac{1}{4} - \frac{1}{9} \right)$$

$$= 1.524 \times 10^6 \text{ m}^{-1}$$

$$\therefore \lambda = 656.3 \text{ nm} \Rightarrow H_{\alpha} \equiv \text{red.}$$

$$\text{If } n = 4 \Rightarrow H_{\beta} = \text{blue, } 486.1 \text{ nm}$$

For $n = \infty$ (the limit of the series).

$$\Rightarrow \lambda = 364.6 \text{ nm (shortest } \lambda \text{ in the series).}$$

Other Series

Lyman, Paschen, Brackett and Pfund

$$\text{-Lyman: } \frac{1}{\lambda} = R \left(\frac{1}{1^2} - \frac{1}{n^2} \right) \quad n = 2, 3, \dots$$

$$\text{-Paschen: } \frac{1}{\lambda} = R \left(\frac{1}{3^2} - \frac{1}{n^2} \right) \quad n = 4, 5, \dots$$

$$\text{-Brackett: } \frac{1}{\lambda} = R \left(\frac{1}{4^2} - \frac{1}{n^2} \right) \quad n = 5, 6, \dots$$

$$\text{-Pfund: } \frac{1}{\lambda} = R \left(\frac{1}{5^2} - \frac{1}{n^2} \right) \quad n = 6, 7, \dots$$

Lyman series wavelengths **ALL U.V**

Balmer series wavelengths **U.V + Visible**

Paschen
 Brackett
 Pfund

}

All I.R